

# Visualizing your Data

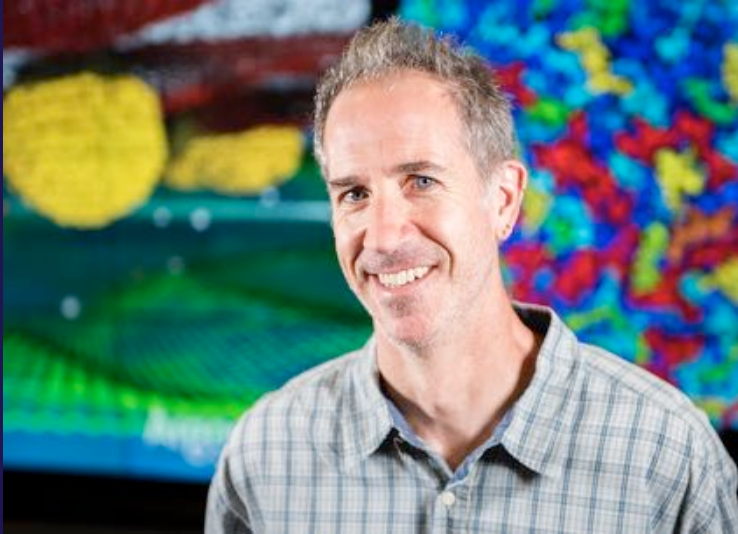
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# ALCF Visualization Group



Joe Insley

Silvio Rizzi



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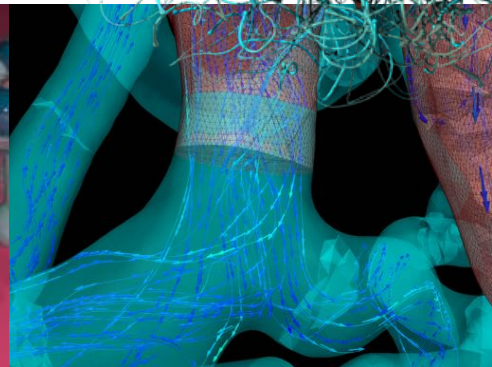
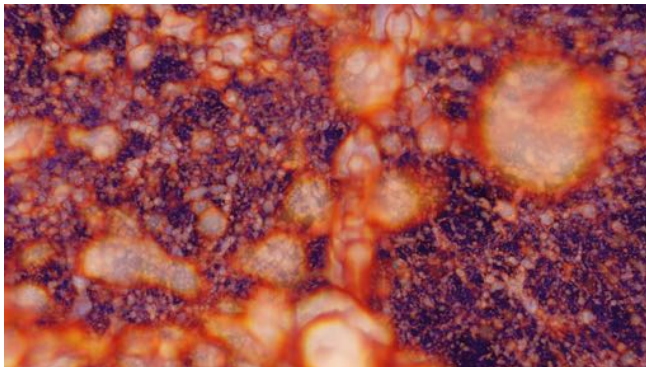
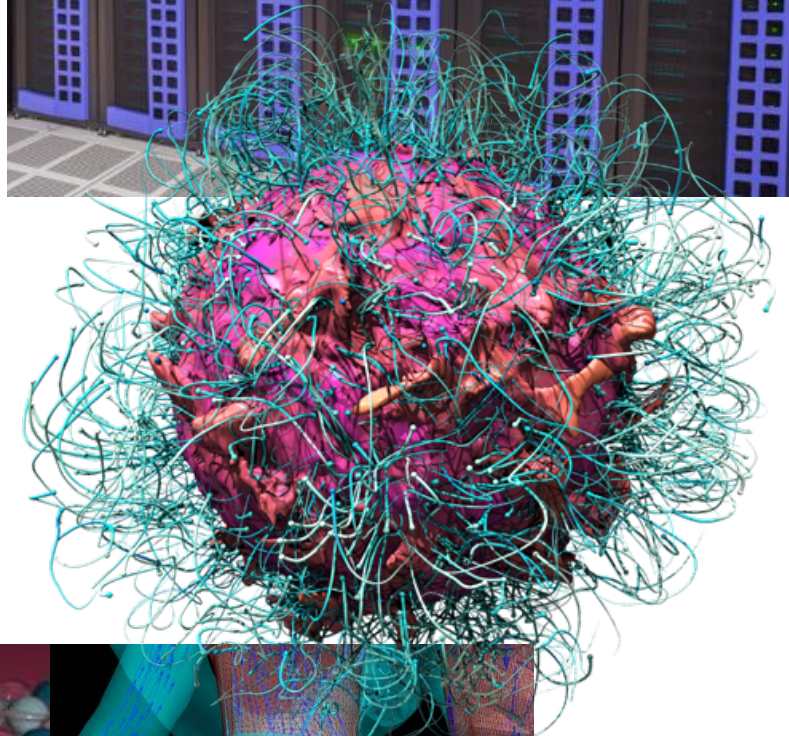
# Here's the plan...

- **Visualization resources, tools and formats**
- **Visualization Success Stories**
  - Blood Flow
  - Cosmology
  - Combustion
- **In Situ Visualization and Analysis**

# Visualization Resources, Tools and Data Formats

# Cooley: Analytics/Visualization cluster

- Peak 223 TF
- 126 nodes; each node has
  - Two Intel Xeon E5-2620 Haswell 2.4 GHz 6-core processors
  - NVIDIA Tesla K80 graphics processing unit (24GB)
  - 384 GB of RAM
- Aggregate RAM of 47 TB
- Aggregate GPU memory of ~3TB
- Cray CS System
- 216 port FDR IB switch with uplinks to our QDR infrastructure
- Mounts the Theta, Eagle, and Grand file systems



# All Sorts of Tools

- Visualization Applications
  - VisIt \*
  - ParaView \*
  - EnSight
- Domain Specific
  - VMD, PyMol, Ovito
- APIs
  - VTK\*: visualization
  - ITK: segmentation & registration
- GPU performance
  - v13: shader-based volume and particle rendering
- Analysis Environments
  - Matlab
  - Parallel R
- Utilities
  - GnuPlot
  - ImageMagick \*

■ Available on Cooley

\* Available on Theta

# ParaView & VisIt vs. vtk

- ParaView & VisIt
  - General purpose visualization applications
  - GUI-based
  - Client / Server model to support remote visualization
  - Scriptable / Extendable
  - Built on top of vtk (largely)
  - In situ* capabilities
- vtk
  - Programming environment / API
  - Additional capabilities, finer control
  - Smaller memory footprint
  - Requires more expertise (build custom applications)



# Data File Formats (ParaView & VisIt)

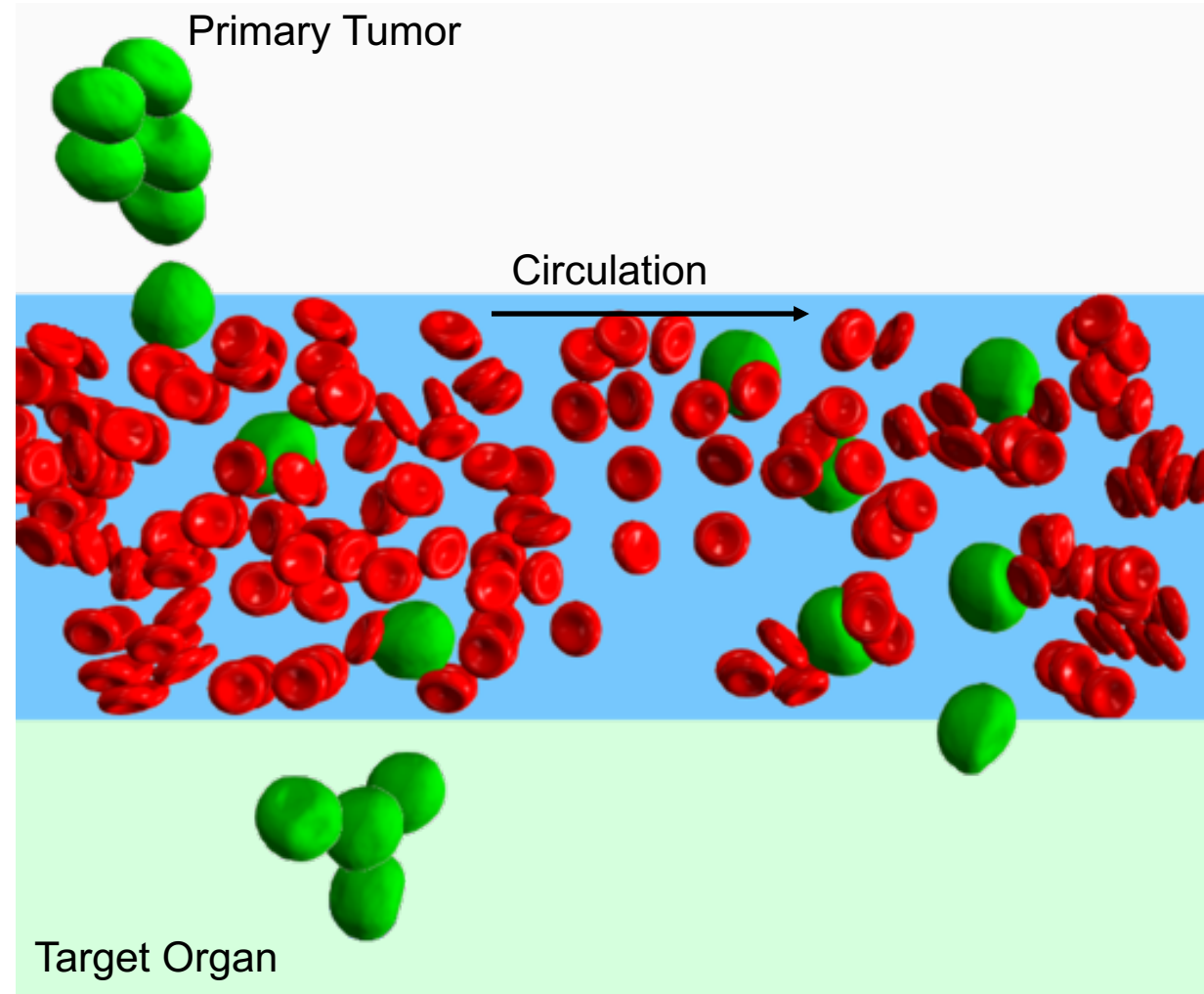
- VTK
- Parallel (partitioned) VTK
- VTK MultiBlock (MultiGroup, Hierarchical, Hierarchical Box)
- Legacy VTK
- Parallel (partitioned) legacy VTK
- EnSight files
- EnSight Master Server
- Exodus
- BYU
- XDMF
- PLOT2D
- PLOT3D
- SpyPlot CTH
- HDF5 raw image data
- DEM
- VRML
- PLY
- Polygonal Protein Data Bank
- XMol Molecule
- Stereo Lithography
- Gaussian Cube
- Raw (binary)
- AVS
- Meta Image
- Facet
- PNG
- SAF
- LS-Dyna
- Nek5000
- OVERFLOW
- paraDIS
- PATRAN
- PFLOTRAN
- Pixie
- PuReMD
- S3D
- SAS
- Tetrad
- UNIC
- VASP
- ZeusMP
- ANALYZE
- BOV
- GMV
- Tecplot
- Vis5D
- Xmdv
- XSF



# Success Stories: Blood Flow

# Motivation

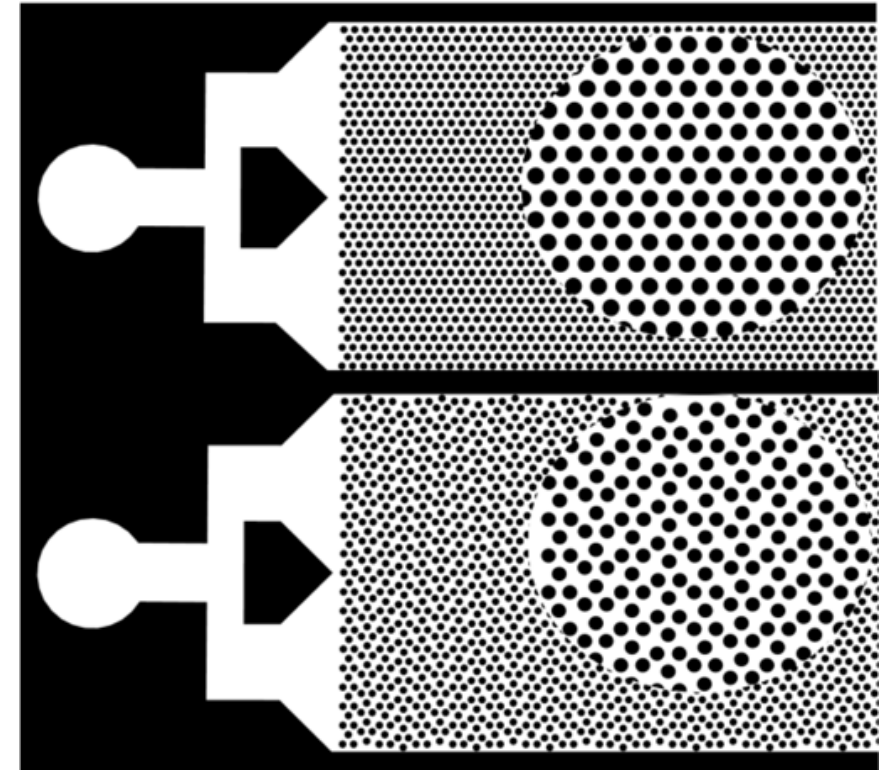
- Circulating tumor cells (CTCs)
  - Can lead to metastasis
  - Metastasis of cancer accounts for 90% of cancer deaths
- Isolation of CTCs
  - Enabling earlier diagnosis
  - Earlier treatment and better outcomes



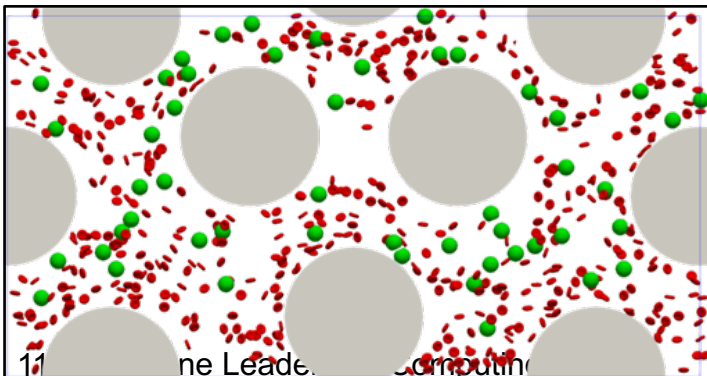
# CTCs in the Device

- CTC physical properties
  - Receptor density on cell surfaces
  - Deformability
  - Size
- Device design and operation parameters
  - Pressure gradient
  - Micropost size and spacing
  - Ligand coating density on microposts
  - Red blood cell (RBC) volume fraction

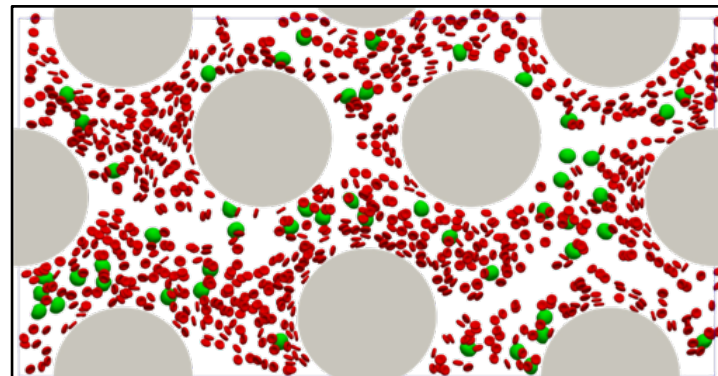
Regular and shifted position of microposts



5% RBC volume fraction

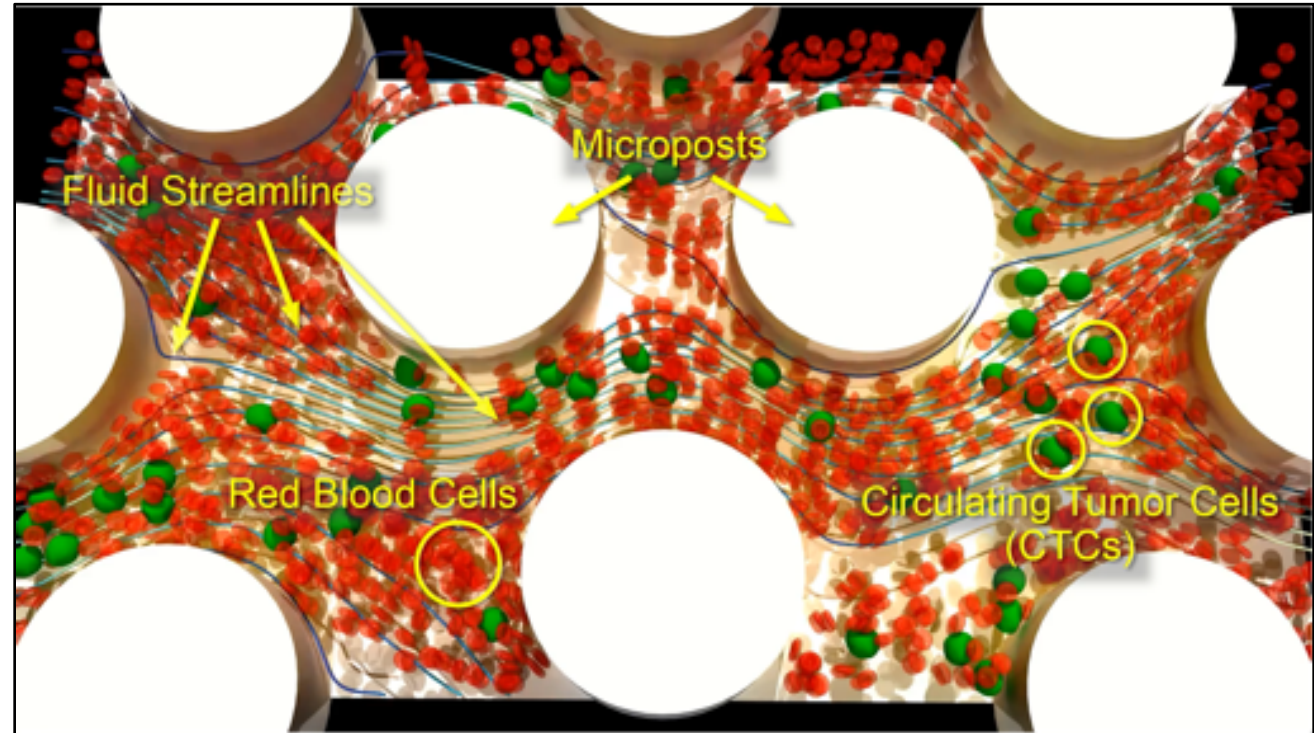


10% RBC volume fraction



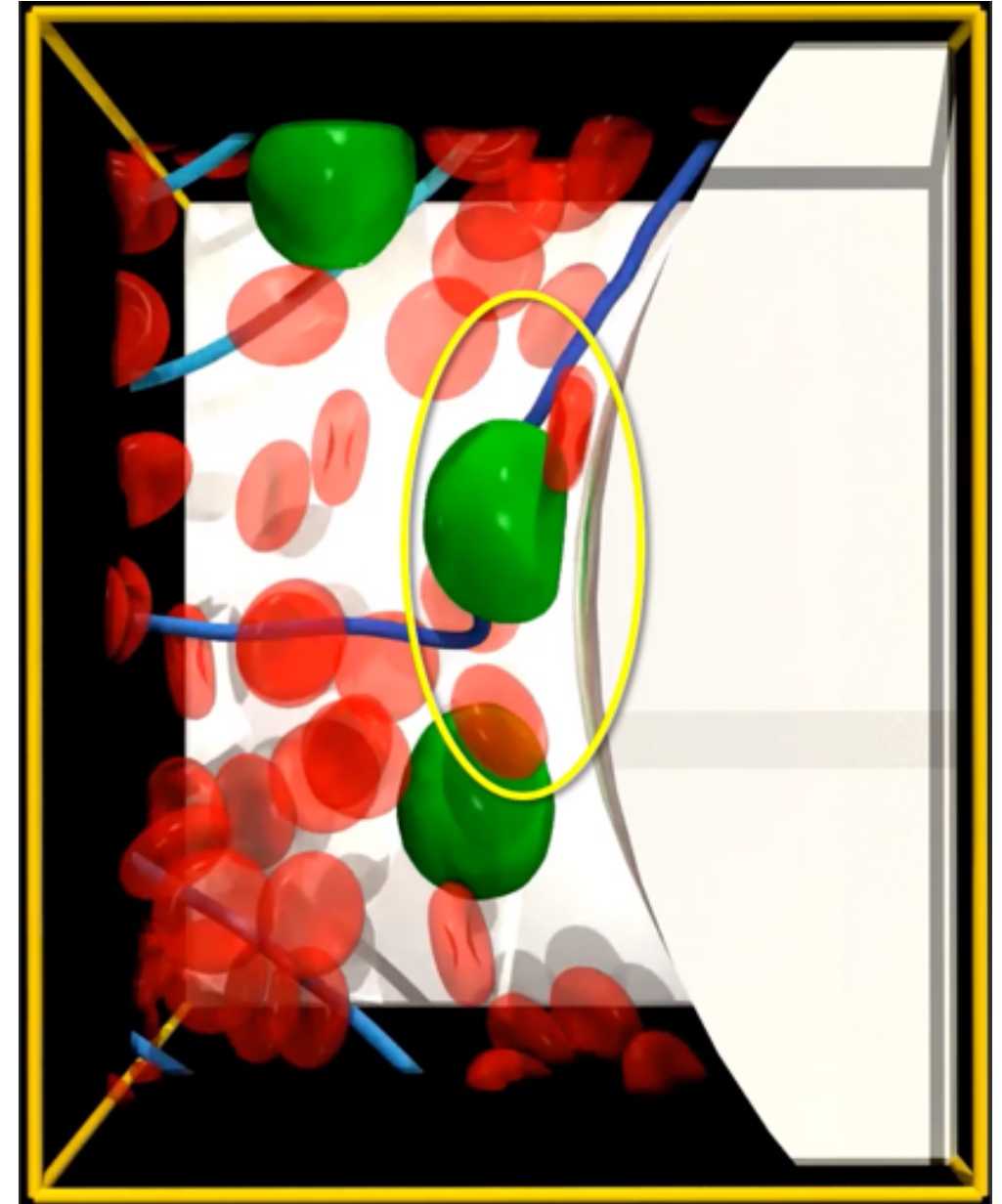
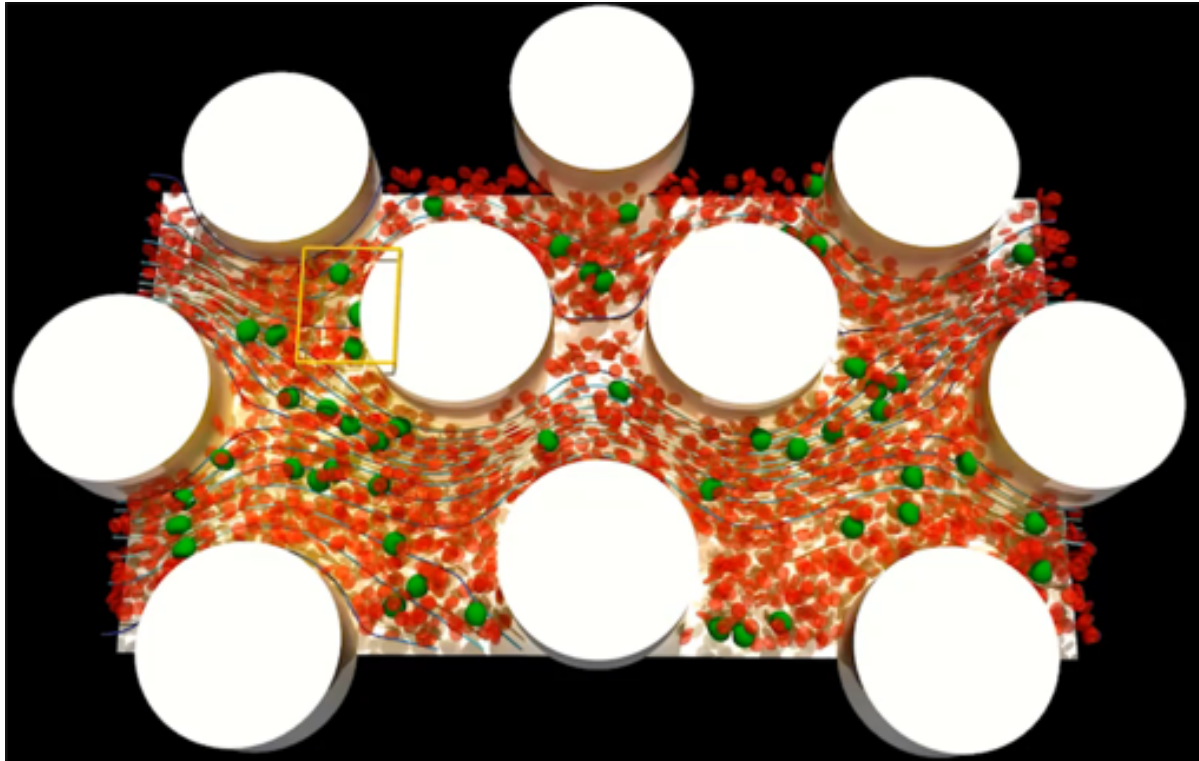
# Modeling Overview

- Fluid solver: Lattice Boltzmann Method (LBM)
  - Highly efficient in parallel processors
  - Palabos software library
- Cell membrane: coarse-grained molecular dynamics
  - Particle based method
  - LAMMPS software library
- Fluid-Solid Interaction: immersed boundary method
  - In-house interface code

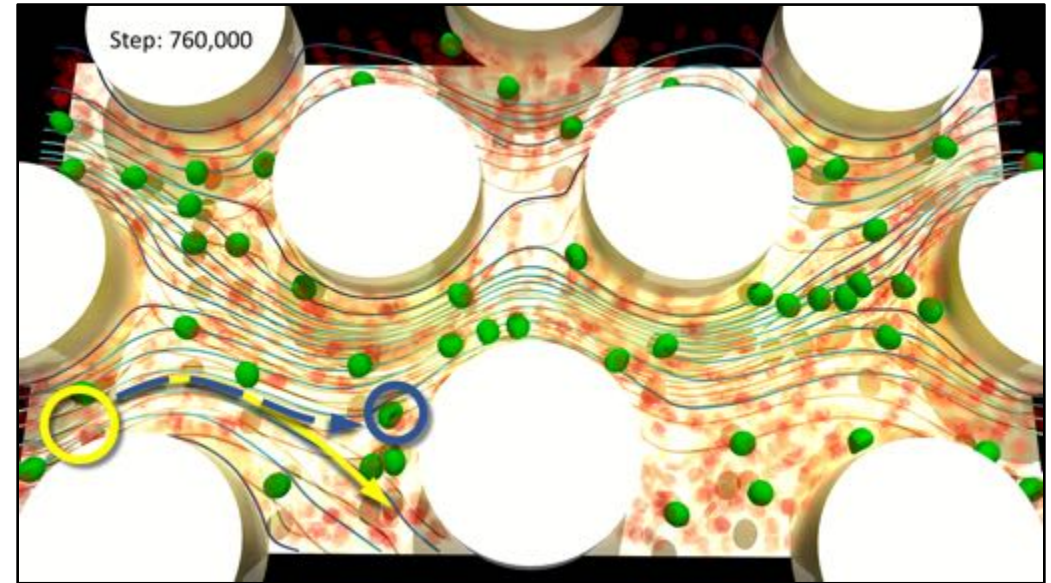
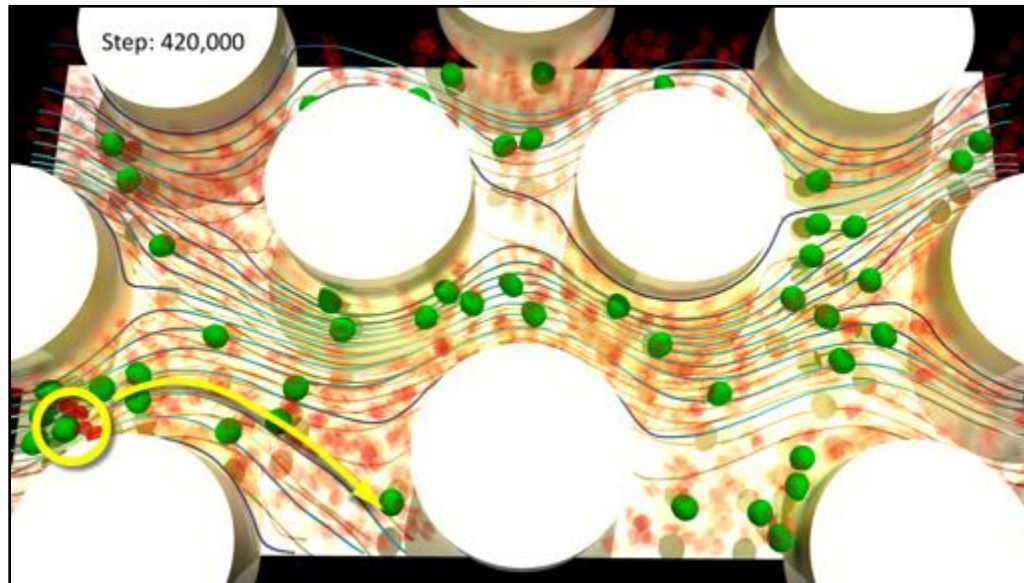
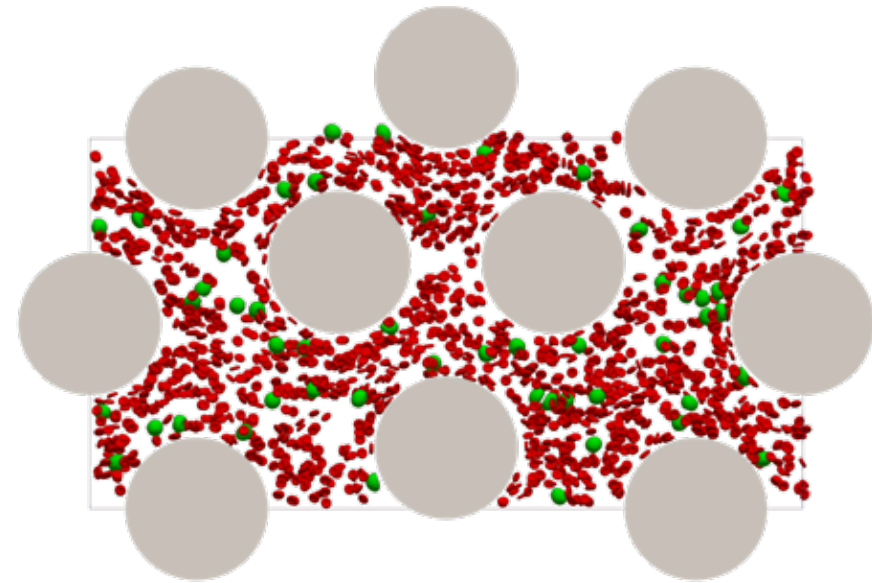
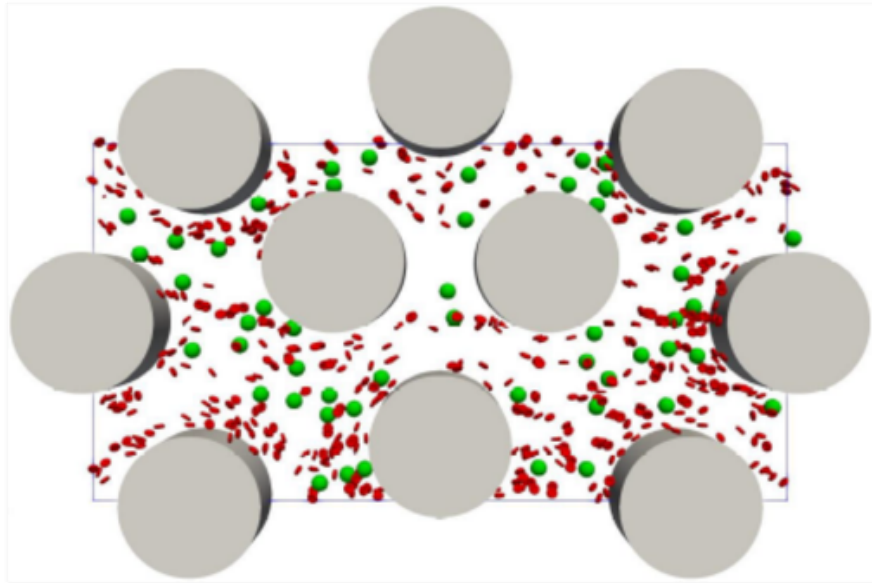


# Adhesion locations

- Impact of
  - CTC size
  - Ligand coating density

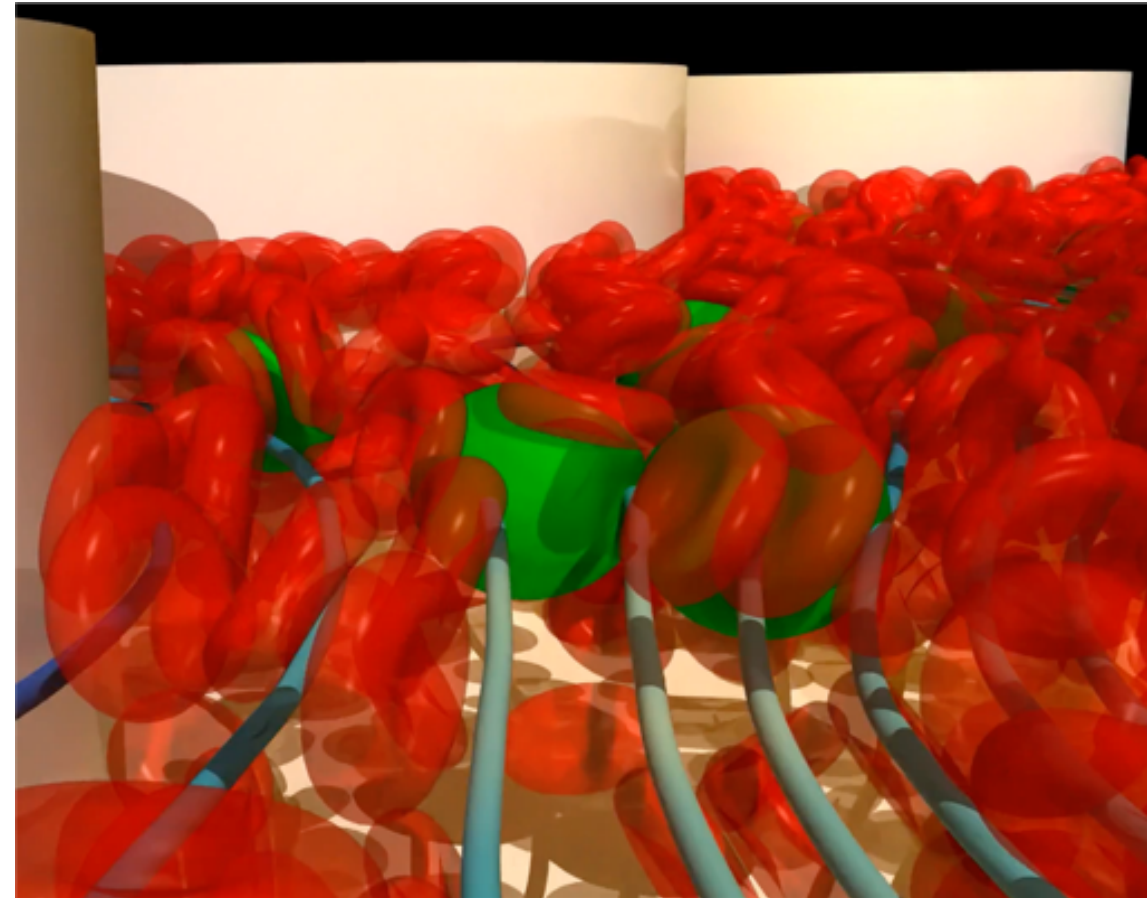


# Transport Trajectories



# Animation

- Simulation
  - Computed on Mira at ALCF
  - 512 processors
  - ~20,000 core hours
  - ~2 Million time steps
- Visualization
  - ParaView
    - Surface rendering of cells
    - Streamline generation of fluid
    - Temporal interpolation of fluid
  - OSPRay



# Visualization of Flow of Circulating Tumor Cells Mixed with Blood Cell Suspensions in Microfluidics

## Science:

**Jifu Tan**

*Northern Illinois University, Department of Mechanical Engineering*

**Michael Hood**

*Northern Illinois University, Department of Mechanical Engineering*

## Visualization:

**Joseph A. Insley**

*Northern Illinois University, School of Art and Design*

*Argonne National Laboratory*

**Michael E. Papka**

*Northern Illinois University, Department of Computer Science*

*Argonne National Laboratory*

**Silvio Rizzi**

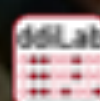
*Argonne National Laboratory*

**Janet Knowles**

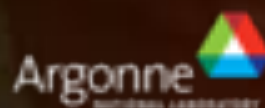
*Argonne National Laboratory*



Northern Illinois  
University



SciLab  
Data Services  
Visualization





# Success Stories: Cosmology

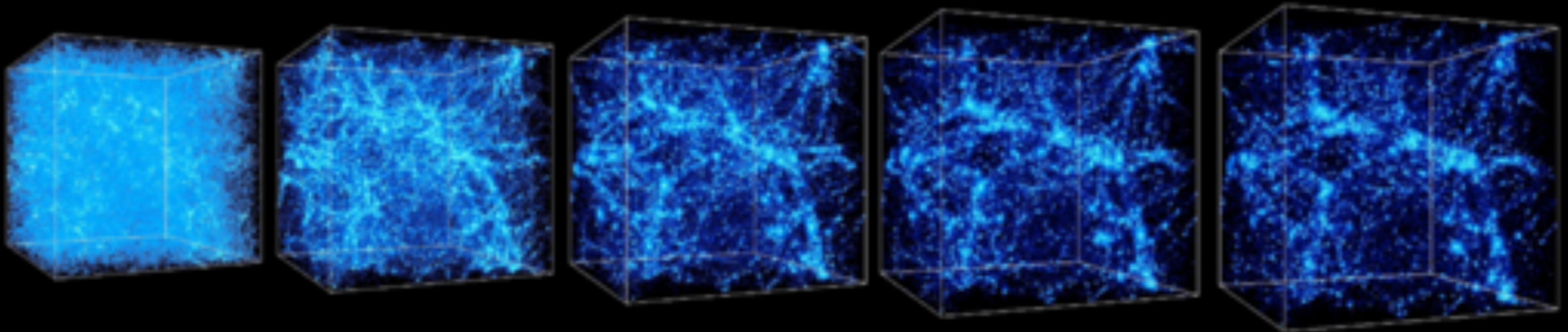
# The Science

Slide courtesy the HACC Team

## Cosmological N-body: Gravity in an Expanding Universe

- Gravity dominates on the largest scales
- Solve Vlasov-Poisson equations
- Use N-body methods with mass tracer particles
  - 6D phase space of positions + velocities
- Large-scale structures condense out of initially smooth distribution of matter

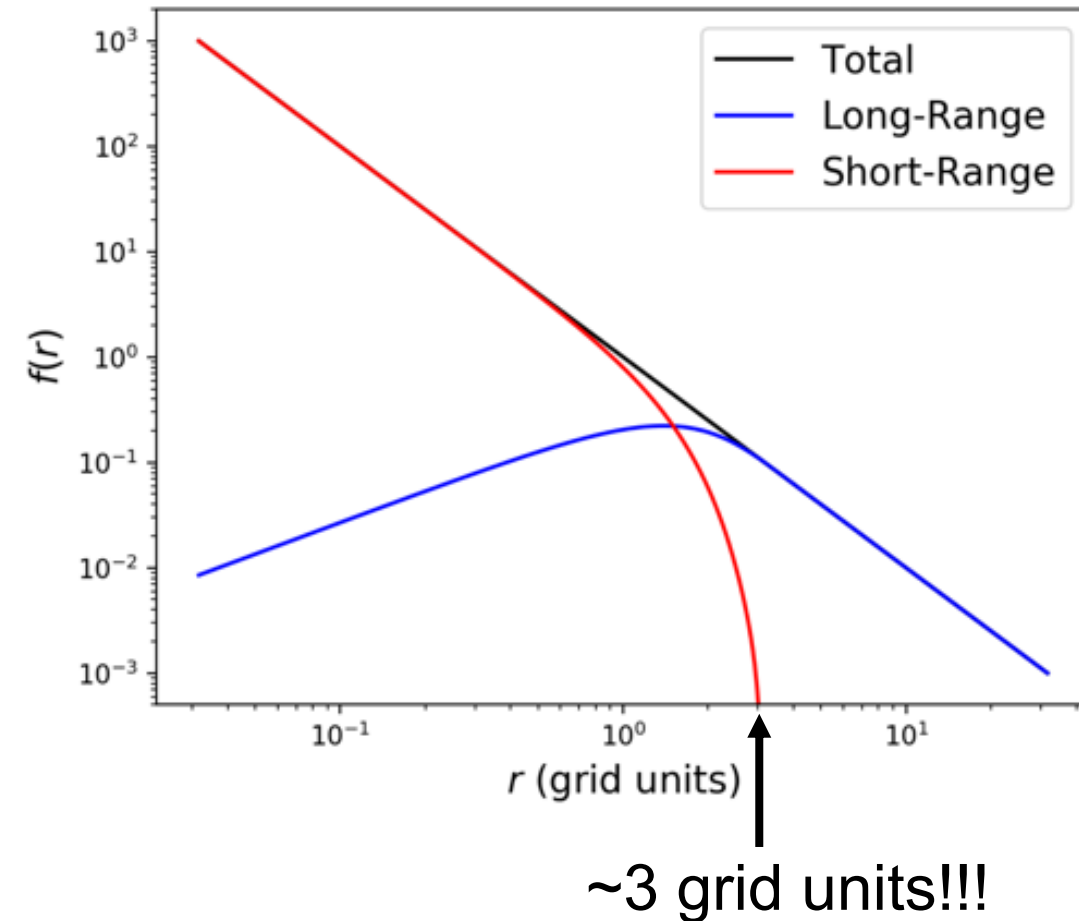
$$\begin{aligned}\frac{\partial f_i}{\partial t} + \dot{\mathbf{x}} \frac{\partial f_i}{\partial \mathbf{x}} - \nabla \phi \frac{\partial f_i}{\partial \mathbf{p}} &= 0, & \mathbf{p} &= a^2 \dot{\mathbf{x}}, \\ \nabla^2 \phi &= 4\pi G a^2 (\rho(\mathbf{x}, t) - \langle \rho_{\text{dm}}(t) \rangle) = 4\pi G a^2 \Omega_{\text{dm}} \delta_{\text{dm}} \rho_{\text{cr}}, \\ \delta_{\text{dm}}(\mathbf{x}, t) &= (\rho_{\text{dm}} - \langle \rho_{\text{dm}} \rangle) / \langle \rho_{\text{dm}} \rangle, \\ \rho_{\text{dm}}(\mathbf{x}, t) &= a^{-3} \sum_i m_i \int d^3 \mathbf{p} f_i(\mathbf{x}, \dot{\mathbf{x}}, t).\end{aligned}$$



# HACC: Gravity Force Splitting

Slide courtesy the HACC Team

- HACC = Hardware/Hybrid Accelerated Cosmology Code
  - Gravity is infinite and unshielded
  - 1 kpc force resolution in 1 Gpc box,  $10^6$  dynamic range
- Operator splitting
  - Kicks: forces used to updated velocities, positions fixed
    - Long-range: Particle-Mesh, deposit onto grid, FFT-based Poisson solver,  $\sim 10^4$  resolution from  $\sim 10k^3$  grids, requires double precision
    - Short-range: Particle-Particle interactions, FLOPS intense, maximize architecture,  $\sim 10^2$  resolution, single precision sufficient
  - Stream: velocities used to update positions, velocities fixed
  - Symplectic integration
- HACC Spectral Force Handover Technology
  - Use low-order Cloud-in-Cell (CIC) deposit
  - Spectral shaping reduces noise and emulates smoother deposit
  - Extremely compact,  $\sim 3$  grid units, limit particle comparisons



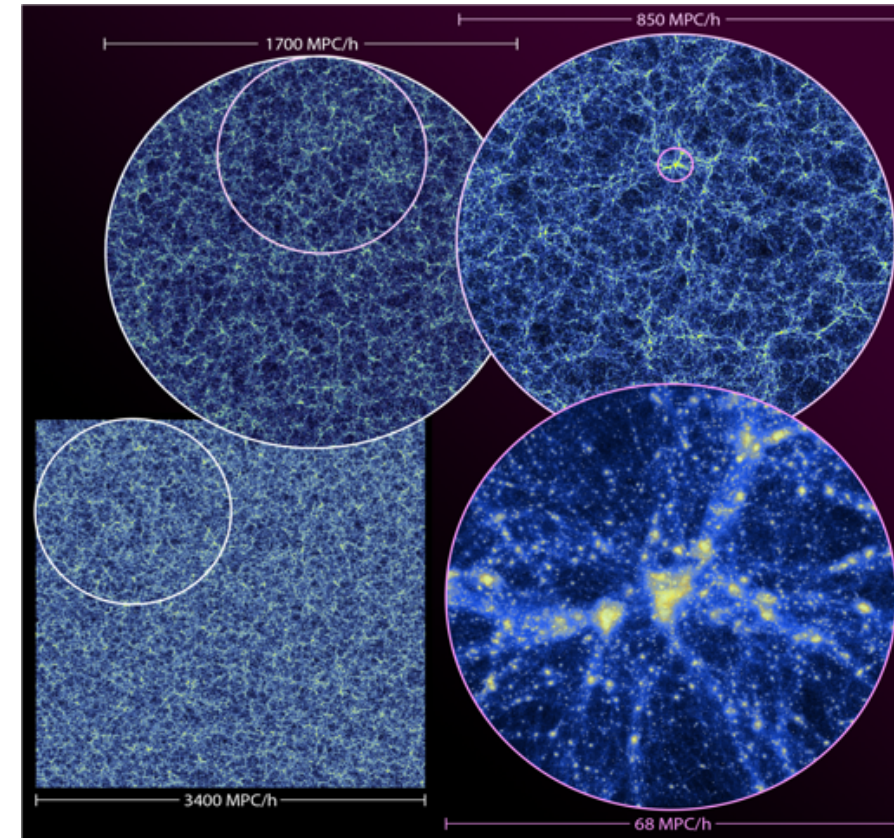
# The Last Journey

## An Extreme-Scale Simulation on the Mira Supercomputer

Slide courtesy the HACC Team

### The Science

- Observational tools: measurements of the growth and distribution of structures on the largest scales, including how galaxies group into clusters and how those clusters are distributed.
- Determine how the content of the universe is divided between normal (baryonic) matter, dark matter, and dark energy, and also to try to understand more about the physical properties of dark matter and dark energy.
- Current and near future astronomical sky surveys can measure positions of millions or billions of galaxies. Massive simulations of cosmological structure formation have become indispensable tools for interpreting the measurements



Zoomed views of a thin density slice of the simulation volume, ending with the largest cluster in the simulation. (Heitmann et al 2021)

# The Last Journey

## An Extreme-Scale Simulation on the Mira Supercomputer

Slide courtesy the HACC Team

### The Impact

- The team used almost 400 million core-hours over the last six months of operation on the Mira supercomputer
- The Hardware/Hybrid Accelerated Cosmology Code (HACC) was used to simulate the movements of 1.24 trillion particles tracing the mass distribution in the universe over cosmic time.
- HACC's CosmoTools infrastructure was used to perform the largest data analysis and reduction operations in-situ, dramatically reducing the amount of data storage required while still producing a rich set of outputs that will be used to support a wide variety of cosmological measurements.
- Outputs from the Last Journey simulation will be used to help with planning and analysis of current and upcoming sky surveys including the Dark Energy Spectroscopic Instrument (DESI), the Legacy Survey of Space and Time (LSST), the Stage 4 Cosmic Microwave Background experiment (CMB-S4), and the SPHEREx space telescope.

Home Mira Activity

	R00	R01	R02	R03	R04	R05	R06	R07	R08	R09	R0A	R0B	R0C	R0D	R0E	R0F
MI																
MC																
	R10	R11	R12	R13	R14	R15	R16	R17	R18	R19	R1A	R1B	R1C	R1D	R1E	R1F
MI																
MC																
	R20	R21	R22	R23	R24	R25	R26	R27	R28	R29	R2A	R2B	R2C	R2D	R2E	R2F
MI																
MC																

Running Jobs

Queued Jobs

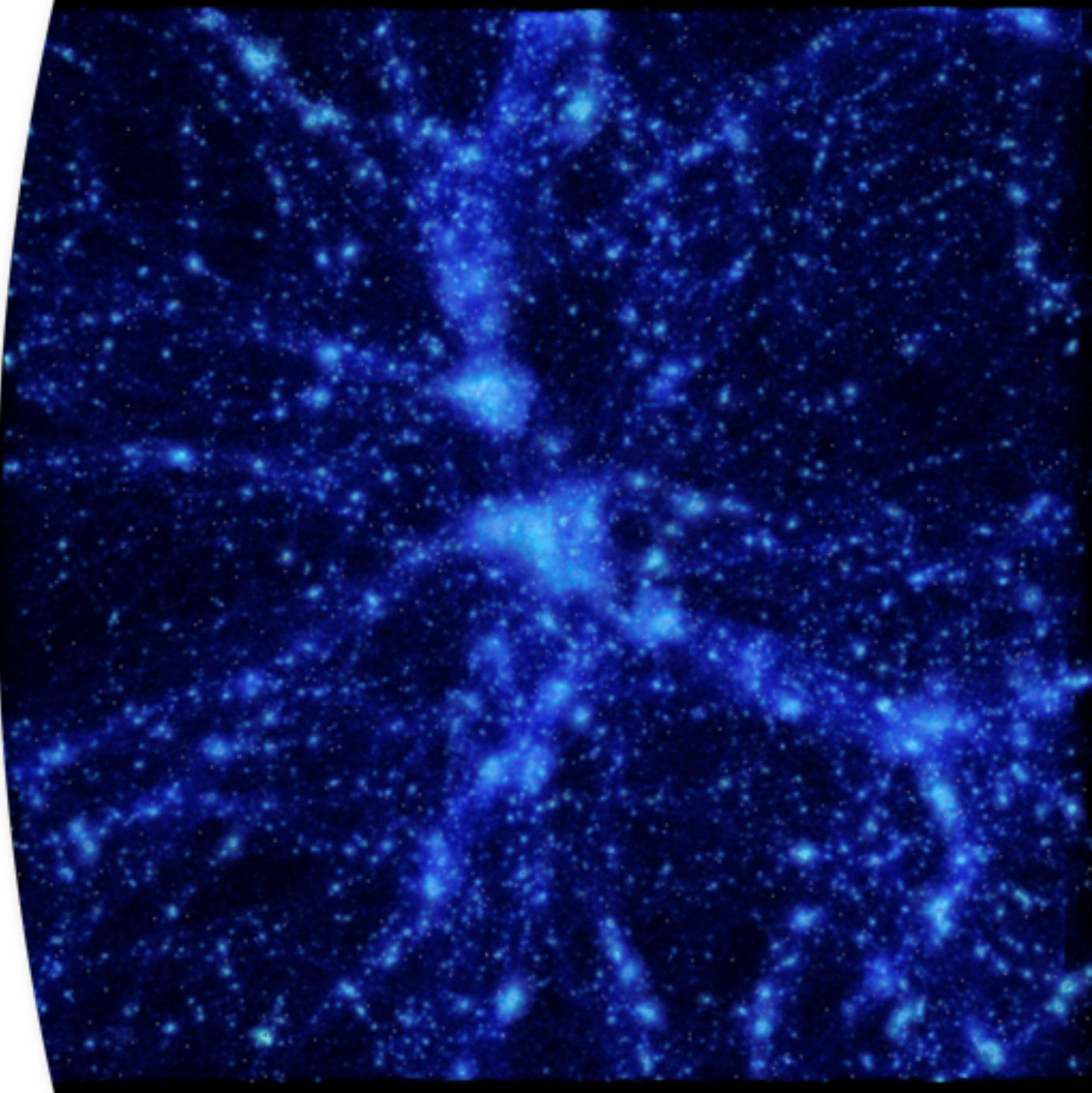
Reservations

Total Running Jobs: 1

Job Id	Project	Run Time	Walltime	Location	Queue	Nodes	Mode
1901442	LastJourney	11:12:29	1d 00:00:00	MIR-00000-78FF1-49152	prod-capability	49152	c2

## Visualizing data from the Last Journey simulation

- Convert data from simulation to regular grids representing density for volume rendering with our vI3 framework
- Convert raw particles and halo particles to VTK for visualization with ParaView
- Combine point rendering and sphere glyphs in ParaView to highlight halos and filament structure



# The Final Extreme-Scale Cosmological Simulation on the Mira Supercomputer: The Last Journey

## Science:

Katrin Heitmann, Nicholas Frontiere, Esteban Rangel, Patricia Larsen, Adrian Pope, Imran Sultan, Thomas Uram, Salman Habib, Hal Finkel, Danila Korytov, Eve Kovacs

## Visualization:

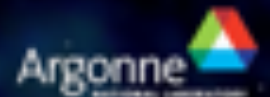
Silvio Rizzi, Janet Y.K. Knowles

Argonne National Laboratory

Joseph A. Insley

Argonne National Laboratory,  
Northern Illinois University

Argonne Leadership  
Computing Facility





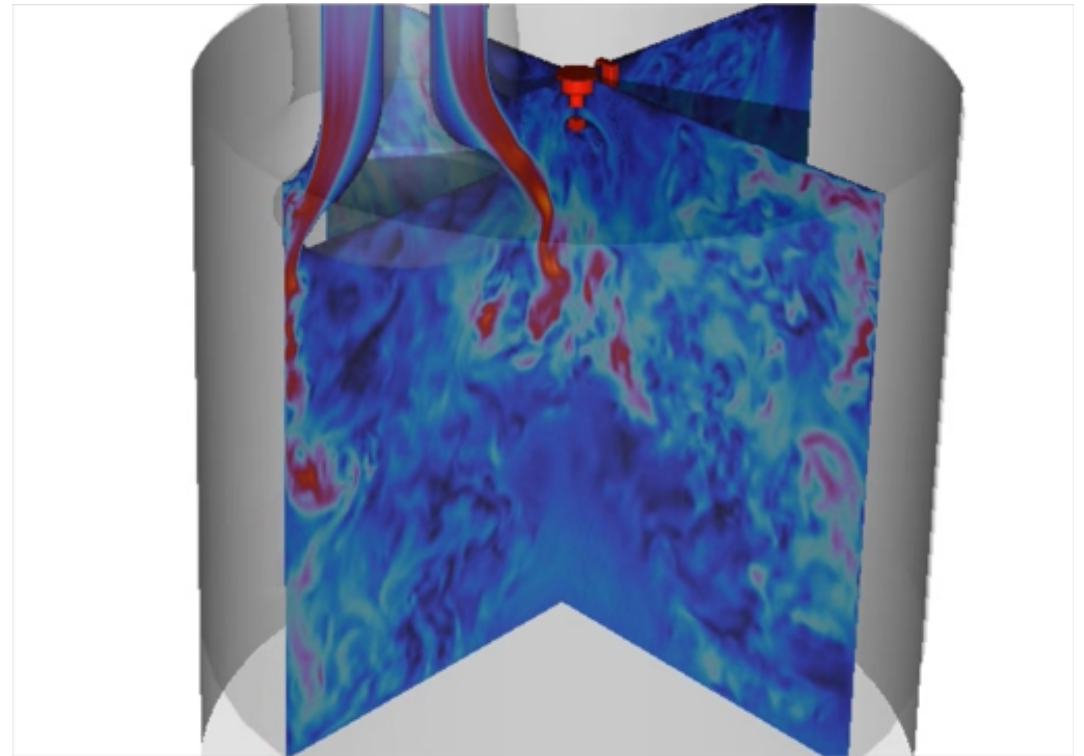
# Success Stories: Combustion

# The Science

## Internal Combustion Engine Simulation



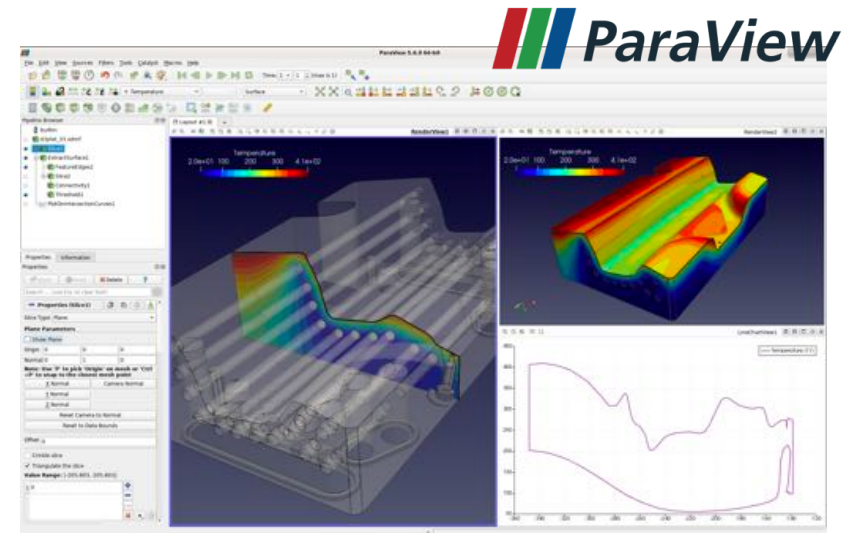
TCC Engine Apparatus



Fluid Dynamics Simulation

# Goal

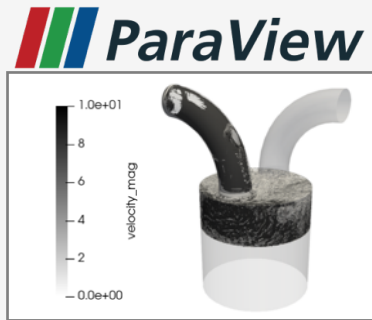
- Provide context to tell the story/explain the science
- Integrate production tools into the existing visualization pipeline
- Tools used:
  - ParaView
  - Maya
  - Substance Painter
  - V-Ray
  - Custom scripts and HPC Resources
  - ffmpeg
  - Premiere/After Effects



# THE VISUALIZATION PIPELINE

## Overview

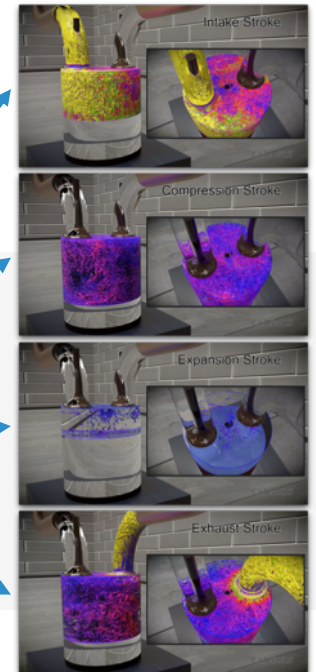
### Visualization Cluster



Export geometry



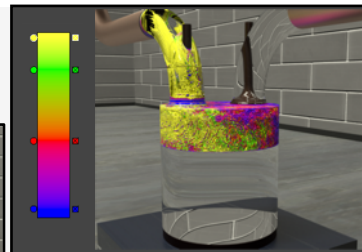
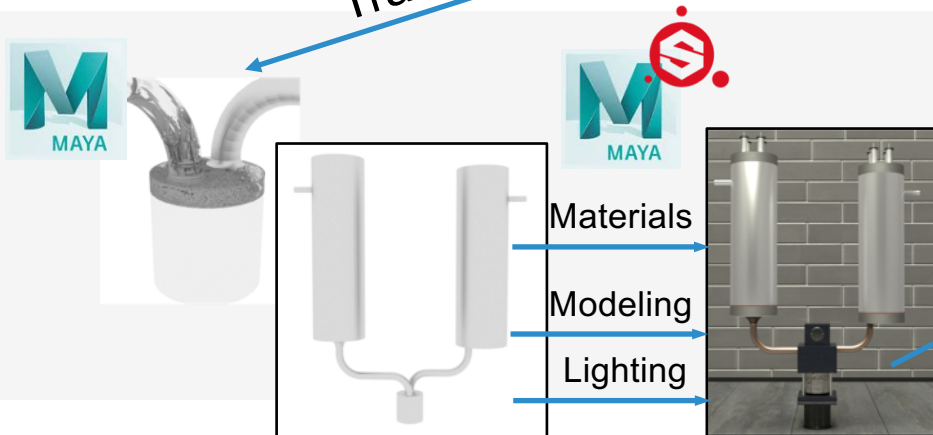
Convert to VRMESH



### Local Workstation

Transfer a few time steps

Transfer VRSCENE



Colors

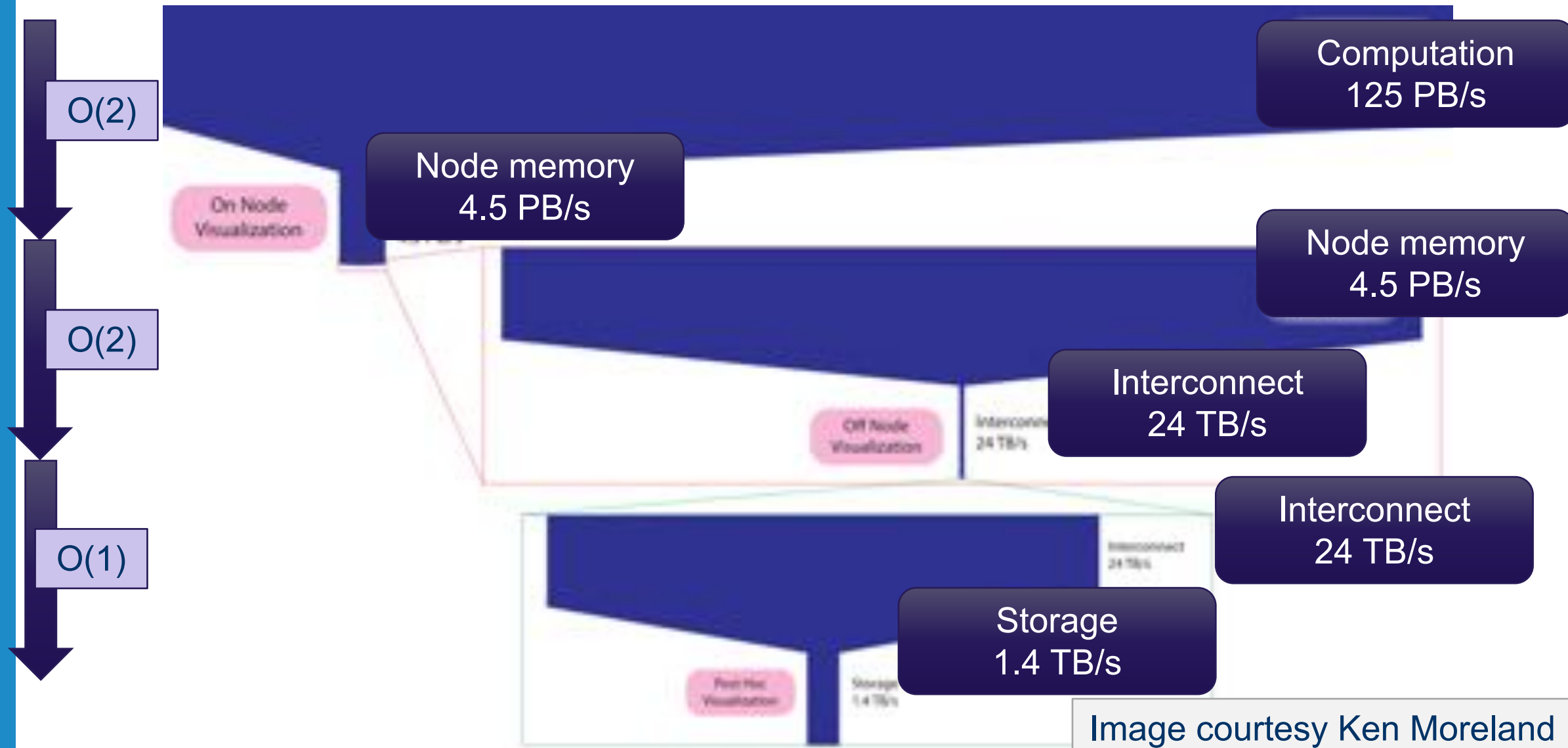
Export to VRSCENE



# Work in Progress: In situ Computational Fluid Dynamics

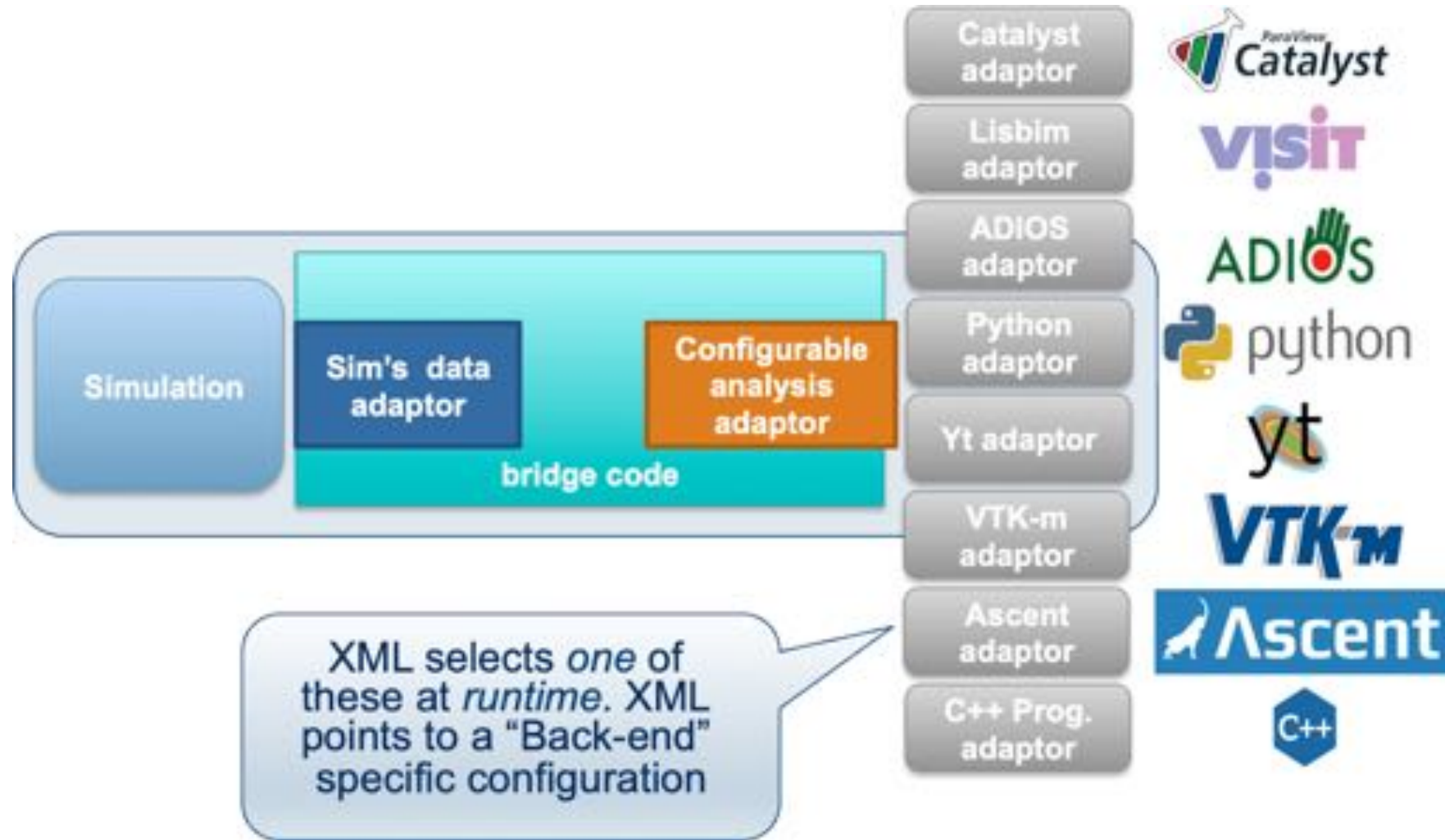
# In Situ Visualization and Analysis: Motivation

Five orders of magnitude between compute and I/O capacity on Titan Cray system at ORNL



# Background

## SENSEI: Scalable Environments for Scientific Explorations In Situ



Loring et al., EGPGV 2020.

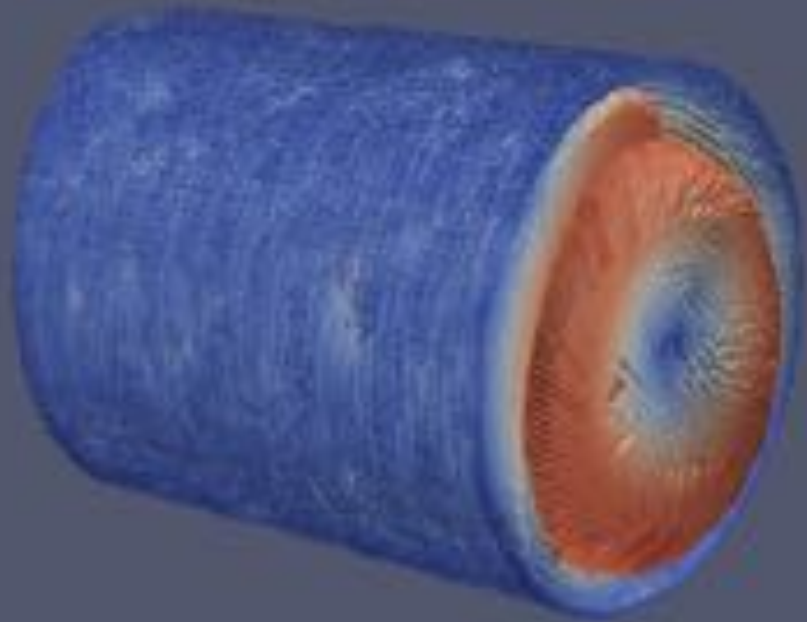
# The Science

Nek5000/NekRS are computational fluid dynamics codes that simulate laminar, transitional, and turbulent incompressible or low Mach-number flows with heat transfer and species transport, which can handle general 2D and 3D domains.

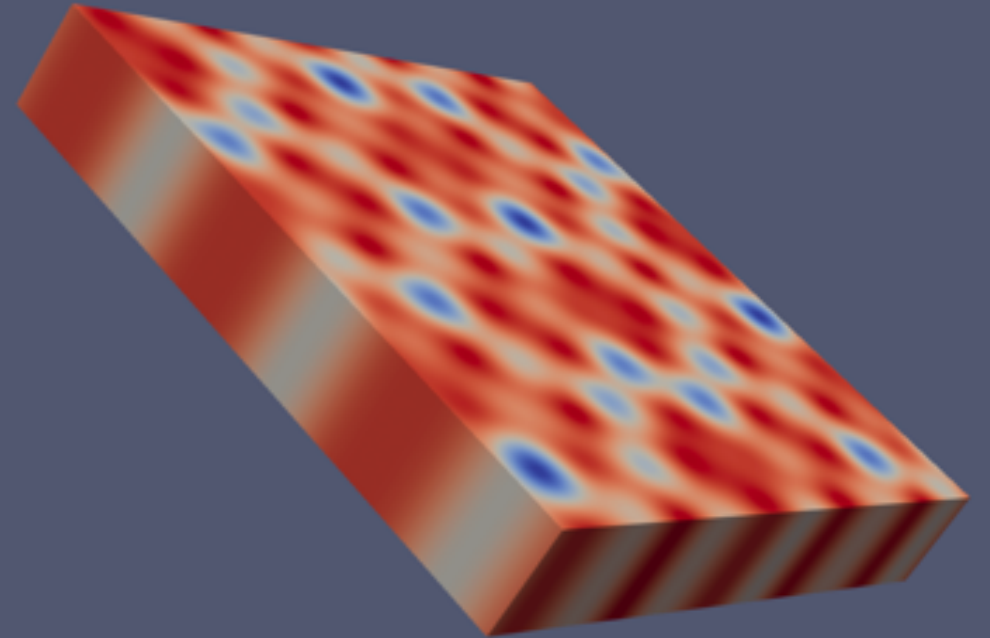
We demonstrated successful full-machine runs (4096 Theta nodes / 262,144 ranks) and generated, at scale, *in situ* renderings, without any data I/O. Additionally, we successfully used ADIOS2 for *in transit* visualization at smaller scales using heterogeneous pipelines. Currently hardening the implementation as well as planning larger runs. The *in transit* runs used 512 ranks on Theta to run the simulation and 96 ranks on Cooley to render it.







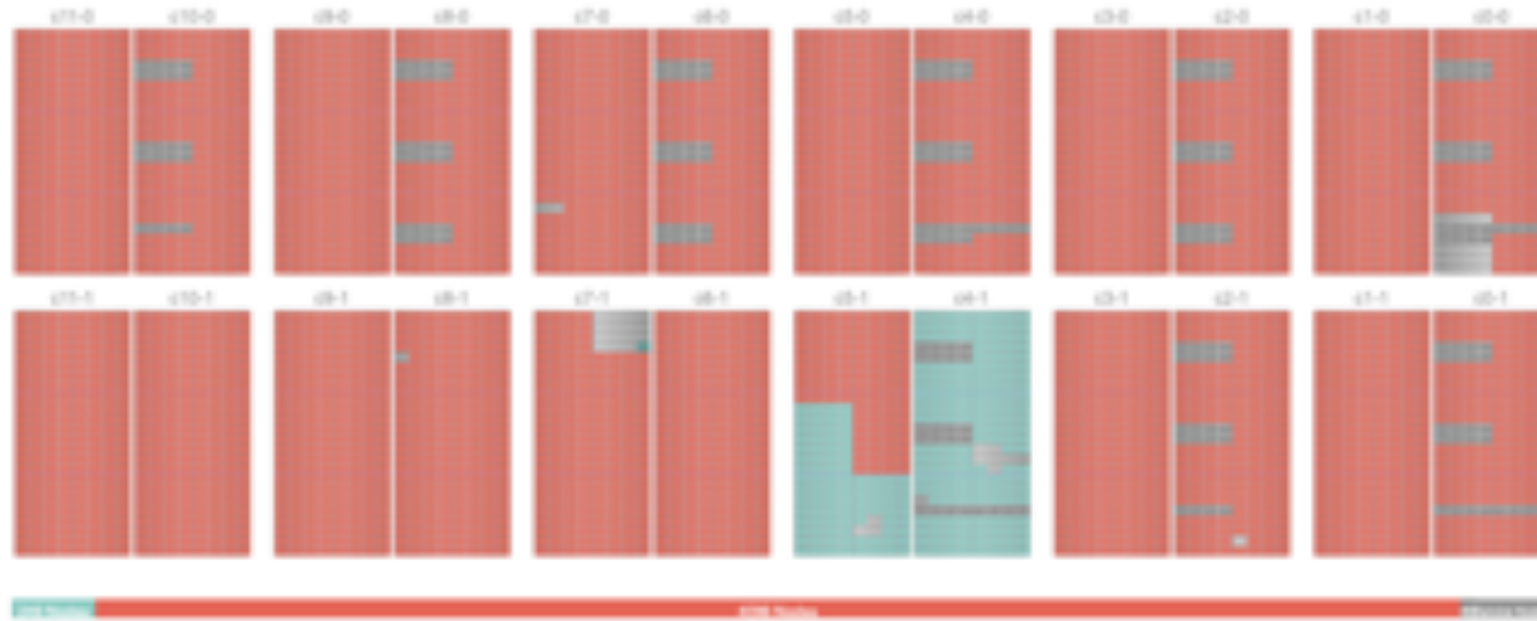
*Swirling, turbulent air during the compression stroke of an internal combustion engine. Data: Saumil Patel*



*NekRS with SENSEI instrumentation*

# Theta Activity

Home Theta 2021 Aug 5 7:14:54 PM Options



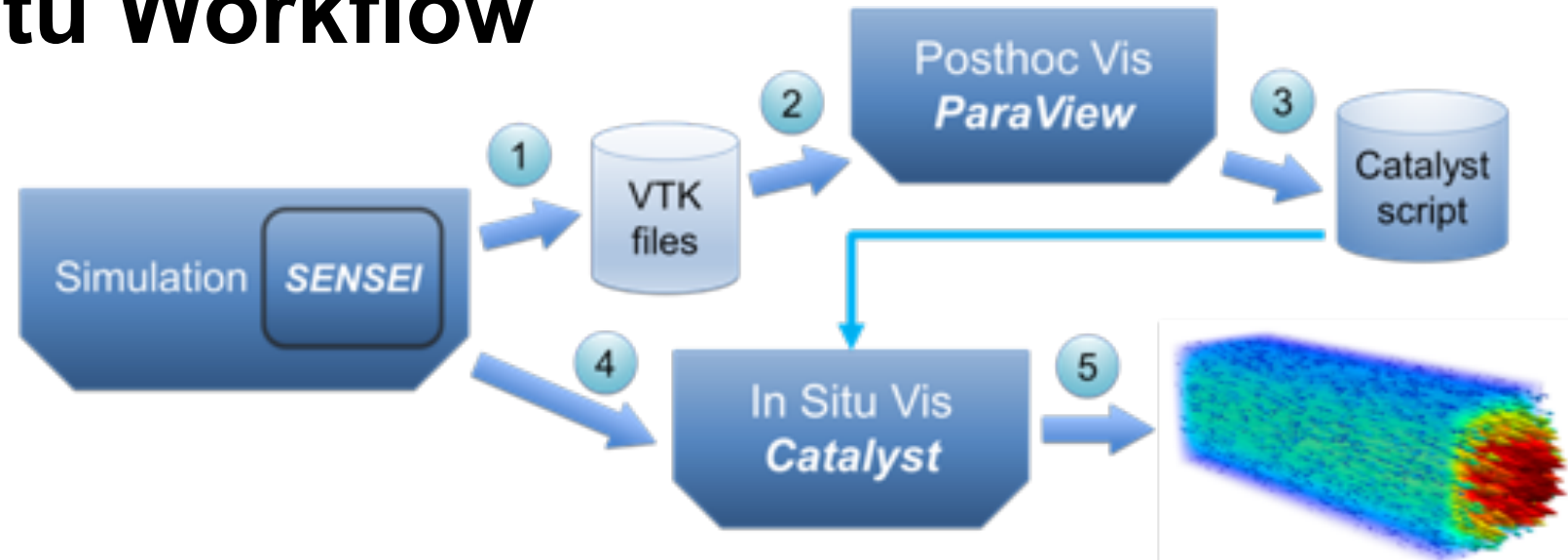
3879 08-10-2021

Running Starting Queued Reservations

Total Running Jobs: 3

Job Id	Project	Nodes	Start Time	Run Time	Walltime	Queue	Mode
538071	SENSEI	4096	7:12:48 PM	00:02:06	01:00:00	default	script
538030	CVD_CityCOVID	248	7:23:12 PM	23:51:42 3d	12:00:00	CVD_Research	script
538073	SENSEI	1	6:44:13 PM	00:30:41	01:00:00	debug-cache-quad	interactive

# Blood Flow In Situ Workflow

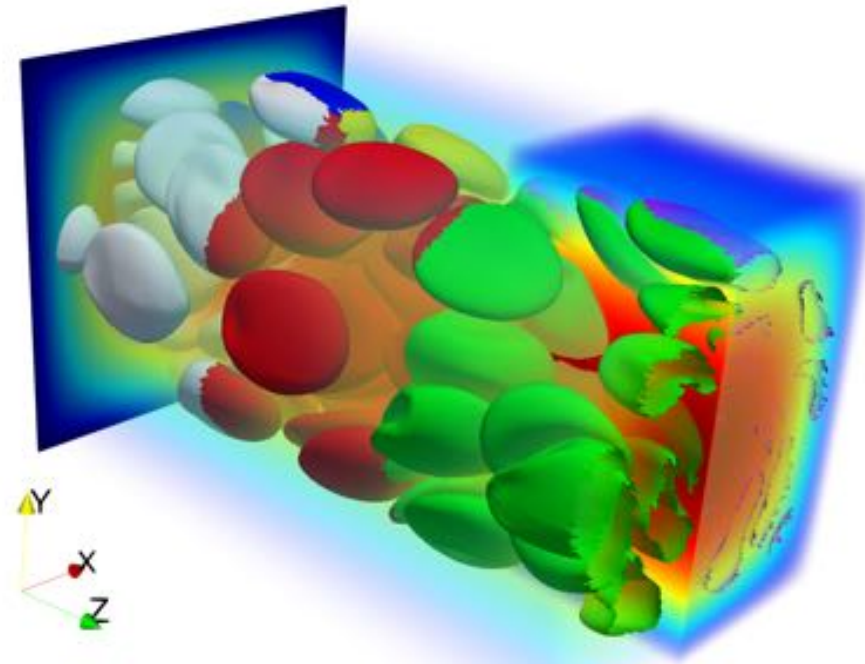


Workflow to create and in situ visualization pipeline<sup>1</sup>

1. Run a small simulation with SENSEI and save a representative data set (PostHocIO).
2. Visualize with ParaView and define desired parameters.
3. Save the state as a Catalyst script
4. Run the simulation with Catalyst enabled and using the Catalyst script.
5. In situ visualization

[1] A. Bucaro et al. "Instrumenting Multiphysics Blood Flow Simulation Codes for In Situ Visualization and Analysis," *2021 IEEE 11th Symposium on Large Data Analysis and Visualization (LDAV)*, pp. 88-89.

# Parallelization Results:

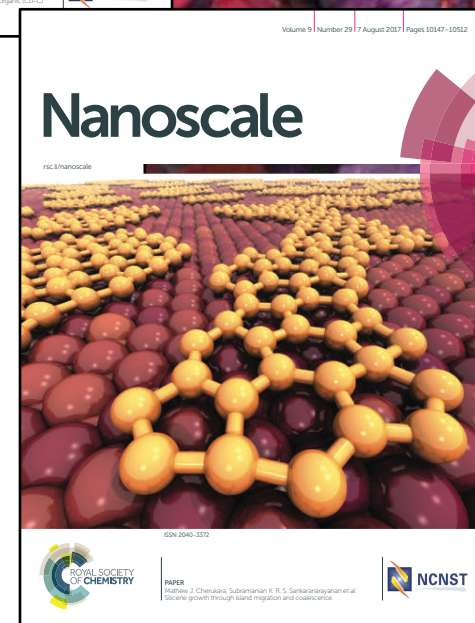
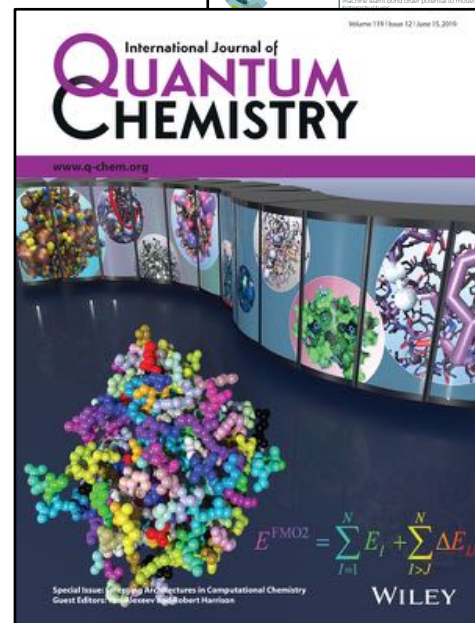
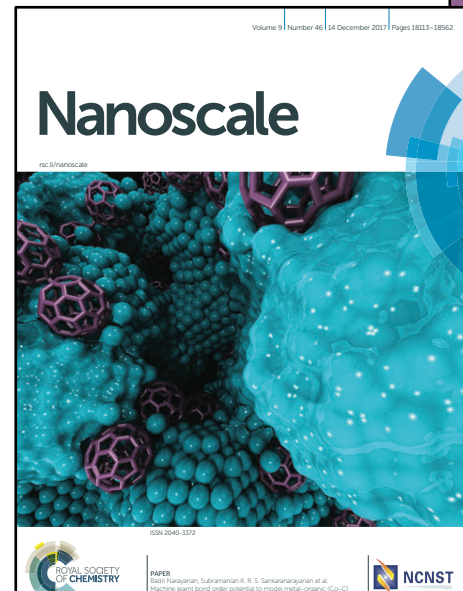


An *in situ* visualization of a blood flow simulation generated using PostHocIO showing how fluid and cells are divided among six processors. Only one 3D subdomain of the fluid velocity is shown in the front for clear visualization.

# *Additional Resources*

# Visualization Help

- [support@alcf.anl.gov](mailto:support@alcf.anl.gov)
- Publication Images & Covers
- Animations
  - SC Visualization Showcase [Best Vis Finalist 2014-2020]
  - APS Division of Fluid Dynamics Gallery of Fluid Motion
- SC Gordon Bell Submissions
- Press Releases
- *InSitu* Vis and Analysis



# QUESTIONS?

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